

1 **Gas emissions at the continental margin west off**  
2 **Svalbard: Mapping, sampling, and quantification**

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25 **Abstract**

26 We mapped, sampled, and quantified gas emissions at the continental margin west  
27 of Svalbard during R/V Heincke cruise He-387 in late summer 2012.

28 Hydroacoustic mapping revealed that gas emissions were not limited to a zone  
29 just above 396 m below sea level (mbsl). Flares from this depth gained significant  
30 attention in the scientific community in recent years because they may be caused  
31 by bottom water-warming induced hydrate dissolution in the course of global  
32 warming and/or by recurring seasonal hydrate formation and decay. We found  
33 that gas emissions occurred widespread between about 80 and 415 mbsl which  
34 indicates that hydrate dissolution might only be one of several triggers for active  
35 hydrocarbon seepage in that area. Gas emissions were remarkably intensive at the  
36 main ridge of the forlandet moraine complex in 80 to 90 m water depths, and may  
37 be related to thawing permafrost.

38 Focused seafloor investigations were performed with the remotely operated  
39 vehicle (ROV) 'Cherokee'. Geochemical analyses of gas bubbles sampled at  
40 about 240 mbsl as well as at the 396-m gas emission sites revealed that the vent  
41 gas is primarily composed of methane (>99.70%) of microbial origin (average  
42  $\delta^{13}\text{C} = -55.7 \text{‰ V-PDB}$ ).

43 Estimates of the regional gas bubble flux from the seafloor to the water column in  
44 the area of possible hydrate decomposition were achieved by combining flare  
45 mapping using multibeam and single beam echosounder data, bubble stream  
46 mapping using a ROV-mounted horizontally-looking sonar, and quantification of  
47 individual bubble streams using ROV imagery and bubble counting. We  
48 estimated that about  $53 * 10^6$  mol methane were annually emitted at the two areas  
49 and allow a large range of uncertainty due to our method ( $9$  to  $118 * 10^6$  mol yr<sup>-1</sup>).  
50 These amounts, first, show that gas emissions at the continental margin west of  
51 Svalbard were in the same order of magnitude as bubble emissions at other  
52 geological settings, and second, may be used to calibrate models predicting  
53 hydrate dissolution at present and in the future, third, may serve as baseline (year  
54 2012) estimate of the bubble flux that will potentially increase in future due to  
55 ever-increasing global-warming induced bottom water-warming and hydrate  
56 dissolution.

57

58 Keywords: vent gas, hydroacoustic flare mapping, hydrate dissociation, global  
59 warming

## 60 **1 Introduction**

61 The Arctic is warming faster than any other region on earth, at the same time, gas  
62 hydrates in Arctic continental margins store significant amounts of methane  
63 (Archer and Buffett, 2005). As hydrates are stable at low temperature and high  
64 pressure conditions, gas hydrates in high-latitude regions that are characterized by  
65 relatively low bottom-water temperatures, can persist in relatively shallow water  
66 depths. Because those regions are highly sensitive to increases in bottom-water  
67 temperatures in the course of global warming shallow hydrates are highly  
68 susceptible to thermal dissociation, which might lead to methane release from the  
69 seafloor. Moreover, methane escaping the seafloor at shallow depths eventually  
70 reaches the atmosphere where it could contribute to the inventory of greenhouse  
71 gases. In that light, findings by Westbrook et al. (2009) were alarming: numerous  
72 gas emissions occurred at the continental margin west of Svalbard concentrated  
73 along a band at seafloor depths just above the 396-m isobath, which is the present  
74 top of the gas hydrate stability zone (GHSZ). During the last three decades the  
75 bottom water at that depth experienced a warming trend of 1°C (Westbrook et al.,  
76 2009). The authors assumed that the warming has induced a deepening of the  
77 upper boundary of the GHSZ from a depth of about 360 m 30 years ago to the  
78 present limit at 396 m, which could have caused hydrate dissociation in the  
79 sediments and, as a consequence, release of gas bubbles. The ‘396-m flares’, as  
80 we call the site here, would be the first site where the hypothesis of global  
81 warming-induced hydrate dissolution may actually be confirmed.

82 Westbrook et al. (2009) offered an alternative hypothesis for the shelf-parallel  
83 occurrences of seafloor gas emissions. Free methane in deep continental slope  
84 sediments may migrate upward along the base of the GHSZ landward to the  
85 depths where it pinches out, which could also explain the clustering of gas  
86 emissions at 396 m depth. A prerequisite of this second hypothesis would be a  
87 capacious gas reservoir in deeper sediments supplying sufficient gas (primarily  
88 methane) to the gas emissions sites. Indeed, data available so far suggest that the  
89 continental margin west of Svalbard is prone to hydrocarbon seepage at the  
90 seafloor: the presence of gas hydrates (below ~600 m water depth) and free gas  
91 below the base of the GHSZ is indicated by the presence of a bottom simulating  
92 reflector (Vanneste et al., 2005; Westbrook et al., 2008; Chabert et al., 2011). In

93 addition, hydrates were recovered from shallow sediments in ~900 m water depth  
94 (Fisher et al., 2011). Gas-related seismic facies occur at the upper slope and outer  
95 shelf (Sarkar et al., 2012; Rajan et al., 2012). Gas emissions occur at the 396-m  
96 flares on the upper slope but also at the outer shelf at water depths up to 150 m  
97 (Westbrook et al., 2009). Typical hydrocarbon seep-related bacterial mats were  
98 observed at the shelf (Knies et al., 2004). Elevated bottom-water methane  
99 concentrations and the stable carbon isotope composition of methane in the water  
100 column indicate seepage at the shelf (Damm et al., 2005; Gentz et al., 2014).

101 A third hypothesis of a seasonally varying thickness of the GHSZ was recently  
102 posed by Berndt et al. (2014). Uranium-Thorium-dating on massive methane-  
103 derived authigenic carbonates sampled at the seafloor at the 396-m flares  
104 ('MASOX site') revealed ages of up to three thousand years. These findings  
105 suggest a long history of methane venting, which argues against the hypothesis of  
106 recent global warming-induced hydrate decay. In addition, seasonal fluctuations  
107 of 1-2 °C in the bottom-water temperature measured with a seafloor-deployed  
108 mooring over a period of almost two years might cause periodic hydrate  
109 formation and dissolution (Berndt et al., 2014). However, a seasonally growing  
110 and declining thickness of the GHSZ should, consequently, result in seasonal  
111 fluctuations in gas bubble emissions, with more intensive emissions during the  
112 time of a retreating GHSZ from about June to December (warmer bottom water)  
113 and less intensive (or no) emissions from January to May (colder bottom water).

114 The amount of hydrate-bound methane that could potentially be released during  
115 dissociation was estimated in several modeling studies at the margin west of  
116 Svalbard but is still uncertain since reported numbers span about three orders of  
117 magnitude. The rates are given as annual amount of mol methane released from  
118 hydrate dissolution per meter of margin segment. The initially reported rate of  
119 global warming-induced release of hydrate-bound methane of  $56.1 \cdot 10^3 \text{ mol yr}^{-1} \text{ m}^{-1}$   
120  $\text{m}^{-1}$  (Westbrook et al., 2009) was later scaled down to  $8.8 \cdot 10^3 \text{ mol yr}^{-1} \text{ m}^{-1}$   
121 (Reagan et al., 2011). For the future, a methane release rate from dissociating  
122 hydrates between  $6.9 \text{ to } 20.6 \cdot 10^3 \text{ mol yr}^{-1} \text{ m}^{-1}$  (10 years) and  $13.2 \text{ to } 72.3 \cdot 10^3$   
123  $\text{mol yr}^{-1} \text{ m}^{-1}$  (30 years) depending on different climate scenarios considered is  
124 expected (Marín-Moreno et al., 2013). Comparably high rates with up to 561 to

125  $935 \cdot 10^3 \text{ mol CH}_4 \text{ yr}^{-1} \text{ m}^{-1}$  kept or released in/from the seasonal gas hydrate mass  
126 were estimated by Berndt et al. (2014).

127 The main objective of this study is to quantify the amount of methane emitted as  
128 gas bubbles from the seafloor to the water column. We assume that most of the  
129 methane flux, is it derived from dissociating hydrate or directly from a free gas  
130 reservoir, is released as gas bubbles. Our study provides a useful mean of  
131 assessing the significance of the bubble flux, it can be used to calibrate models of  
132 hydrate dissolution, and, further, it can serve as base-line (year 2012) estimate of  
133 the methane flux that is likely to increase in future due to the ongoing warming  
134 trend. The quantification is based on the combination of ship-borne systematic  
135 hydroacoustic flare mapping and ROV-based estimation of the bubble flux of  
136 individual bubble streams. A further objective of our study is to map the  
137 distribution of gas emissions at the shelf and the upper continental slope west of  
138 Svalbard. Although we are not able to contribute to the ongoing discussion  
139 whether or not hydrate dissolution is the cause for the bubble emissions, flare  
140 distributions determined in the study area put the significance of the 396-m flares  
141 into perspective. Finally, samples of gas bubbles and geochemical analyses give  
142 insight into the genesis (thermogenic versus microbial) of emitted gas.

143

## 144 **2 Study Area**

145 The study area is located west of Svalbard (Fig. 1). The continental margin was  
146 shaped by the advances and retreats of the ice sheet covering Svalbard and the  
147 Barents Sea during the Pliocene-Pleistocene (Solheim et al., 1998; Vorren et al.,  
148 1998). Fast-flowing ice streams created the cross-shelf troughs seaward of the  
149 major fjord systems Kongsfjord and Isfjord. The inter-trough region west of Prins  
150 Karls Forland was covered by slow-flowing ice sheets with the shelf break  
151 marking approximately the seaward extent of the maximal ice coverage (Landvik  
152 et al., 1998). The shelf was flooded as glacial ice retreated about 13000 years ago  
153 (Landvik et al., 2005). Large areas of the shelf were mapped by the Norwegian  
154 Hydrographic Survey (Landvik et al., 2005) and the University of Tromsø  
155 (Ottesen et al., 2007). The existing multibeam data cover the shelf area east and  
156 north of the area shown in Figure 2 with some overlap in the central part. The

157 forlandet morain complex is a pronounced ridge system at the middle slope with a  
158 crest in about 90 m water depth (Landvik et al., 2005). During a cruise in 2011  
159 with the R/V James Clarke Ross gas emissions were found at the forlandet morain  
160 complex (Wright, 2012), an area that for simplicity we call Area 1 in the  
161 following. Additional evidence for hydrocarbon seepage at the shelf was  
162 presented by Knies et al. (2004) who discovered seep-typical sulfur-oxidizing  
163 bacterial mats using ROV.

164 The gas emissions discovered by Westbrook et al. (2009) are located at the outer  
165 shelf (Area 2 in this study) and upper continental slope (Area 3). The  
166 misalignment between gas vents at ~240 m water depth (Area 2) and at 396 m  
167 (Area 3) is caused by the combined action of a slump (Fig. 2) acting as seal for  
168 upward migrating fluids and glacigenic debris flows, which channel fluids along  
169 their base landward, as geophysical studies revealed (Rajan et al., 2012; Sarkar et  
170 al., 2012). Further landward of the prograding glacigenic sequences, pockmarks  
171 exist at the seafloor (Fig. 2) and a seismic image shows that one pockmark was  
172 underlain by an acoustic pipe structure but as no gas emissions were observed so  
173 far, they are probably relict structures of fluid emission (Rajan et al., 2012).

174 Two high-resolution seismic studies were carried out in the area of potential  
175 global-warming induced hydrate dissociation (Area 2 and 3) that led to different  
176 conclusions. The study by Rajan et al. (2012) focused on the region including  
177 Area 2 and the northernmost part of Area 3 (Fig. 2) that are affected by glacigenic  
178 debris flows. The authors imaged a gas cloud in the sediment below the landward  
179 limit of the GHSZ that they interpret as possible migration pathway of deep  
180 (thermogenic) gas. They conclude that the gas may be temporarily sequestered as  
181 gas hydrates but seismic evidence for this is lacking and, thus, any involvement of  
182 global-warming induced hydrate dissociation is speculative. However, based on a  
183 seismic data set covering the entire Area 3, Sarkar et al. (2012) argue, that  
184 evidence for fault-controlled gas migration from deeply-buried sediments, which  
185 could explain the contour-following trend of the flares originating at 396 m water  
186 depth is missing. Instead bright spots at shallow sediment depths close to the  
187 landward limit of the GHSZ, would be in accordance with global-warming  
188 induced hydrate dissolution.

189 While glacial sedimentation was predominant at the shelf and upper slope, the  
190 distal slope was influenced by hemipelagic sedimentation and bottom water  
191 currents, leading to the development of contourite drifts (Eiken and Hinz, 1993).  
192 Vestnesa Ridge is a contourite with evidence for a very active hydrocarbon  
193 venting system (Hustoft et al., 2009). Southeast of Vestnesa Ridge in Area 4  
194 pockmark-like seafloor depressions exist between 800 and 1200 m water depth  
195 (Fig. 1). The presence of gas hydrates in the sediments was inferred from a well  
196 pronounced bottom simulating reflector (Sarkar et al., 2012) and proven by  
197 gravity coring (Fisher et al., 2011).

198

### 199 **3 Material and Methods**

200 The study is based on R/V Heincke cruise No. 387 (20 Aug to 9 Sept 2012)  
201 conducting research in the area west of Svalbard (Sahling et al., 2012). The  
202 multibeam echosounder Kongsberg Maritime EM 710 was employed for seafloor  
203 charting and water-column flare mapping. The system operates at frequencies  
204 between 70 and 100 kHz. It has 200 beams each with an opening angle of 1°  
205 across track and 2° along track. The footprint of the echosounder across track is  
206 therefore about 1.7% of the water depth. Two data sets for seafloor mapping (\*.all  
207 files) and water column mapping (\*.wcd files) were recorded (available online:  
208 <http://doi.pangaea.de/10.1594/PANGAEA.816220>). Seafloor data was processed  
209 with MB Systems (Caress and Chayes, 2001) and water column data with the  
210 program package by the company Quality Positioning Services BV (QPS)  
211 including FM Midwater and Fledermaus. Four sound velocity profiles were  
212 obtained during the cruise using a MIDAS sound velocity probe (company  
213 Valeport).

214 Scientific single beam echosounder EK 60 operates with up to four frequencies  
215 but for the purpose of this study, only the 38 kHz frequency was analyzed for  
216 mapping and flare classification purposes. Data were recorded with the ER 60  
217 software, stored as \*.raw files (available online:  
218 <http://doi.pangaea.de/10.1594/PANGAEA.816056>), and processed using the  
219 readEKRaw MATLAB toolkit (by Rick Towler, NOAA Alaska Fisheries Science  
220 Center; available online: <http://hydroacoustics.net/viewtopic.php?f=36&t=131>).

221 The toolkit was used to convert the data into Sv, which is the volume  
222 backscattering per unit volume expressed in dB re 1 m<sup>-1</sup>. Sv is often used when  
223 individual targets are very small in the sampled volume as several echoes are  
224 combined to give a certain signal level. A toolkit for mapping flares was  
225 designed. This consists of an interface where the user is reading echosounder  
226 traces and is asked to pick manually the flares that appear. For each selected flare,  
227 an Id (with the format DayMonthNumbering) is given and its characteristics are  
228 stored (Supplementary material S1): the date and time at which it was observed,  
229 its longitude and latitude, its strength as the weighted sum of all Sv levels within  
230 its trace area, and finally its height. The weighted sum of all Sv levels was made  
231 on a linear scale with the purpose of classifying flares into strong and weak.  
232 Locations of flares were plotted with GMT using color coding for classifying  
233 strong and weak flares (threshold arbitrarily set at 4 dB re 1 m<sup>-1</sup>; Fig. 2).

234 The remotely operated vehicle (ROV) MARUM-Cherokee is a mid-size  
235 inspection class vehicle manufactured by Sub-Atlantic, Aberdeen. Underwater  
236 positioning was obtained using the ultra-short baseline system GAPS by Ixsea.  
237 Scientific payload of the ROV was a modified, small-sized version of the  
238 pressure-tight Gas Bubble Sampler (GBS; Pape et al., 2010), custom made bubble  
239 catchers, and horizontally scanning sonars (Imagenex 881A or Tritech) mounted  
240 on top of the vehicle to allow 360° sonar view. Still images were acquired with a  
241 5 megapixel Kongsberg OE-14 camera. Videos were recorded with a Tritech  
242 Typhoon PAL camera and stored electronically in AVI format.

243 The volume flux of bubbles was estimated using a bubble catcher and visually  
244 using the video. Scaling of the images was obtained by placing objects of known  
245 dimensions (such as the ROV-manipulator) into the plane where the bubbles  
246 occur. Due to the low shutter speed, bubbles appear blurred as long ellipsoids in  
247 the video frames and, therefore, only one bubble diameter could be measured.  
248 From each measure, volumes were calculated assuming spherical bubbles and  
249 fluxes were inferred by multiplying the average bubble volume with the emission  
250 frequency. The volume flux was then converted to mass flux assuming that the  
251 gas consists of pure methane and considering the compressibility of methane  
252 (compressibility = 0.91 at 380 m water depth, 39 bar, 4 °C; compressibility = 0.93  
253 at 240 m water depth, 25 bar, 4 °C). A SBE911plus Sea-Bird Electronic CTD was

254 used to acquire hydrographic parameters. Gas collected with the GBS was  
255 analyzed with a two-channel 6890 N (Agilent Technologies) gas chromatograph  
256 described in detail in Pape et al. (2010). Hydrate phase boundaries were  
257 calculated using the HWHYD U.K. software (Masoudi and Tohidi, 2005).

258

## 259 **4 Results**

### 260 **4.1 Flare mapping**

261 A total of 1920 nautical miles of hydroacoustic profiles were acquired during the  
262 He-387 cruise (Fig. 1). For simplicity, we subdivided the region in five areas.  
263 Flares in the water column were found at the continental shelf (Area 1), close to  
264 the shelf break (Area 2), and at the upper continental slope (Area 3), but not  
265 above the pockmarks (Area 4), and along the 396-m depth contour further north  
266 (Area 5).

267 Numerous flares occurred at the shelf and upper slope west of Prins Karls Forland  
268 (Fig. 2). Gas emissions concentrate in Areas 1, 2, and 3. Emission sites in Areas 2  
269 and 3 correspond to those discovered by Westbrook et al. (2009) at water depth  
270 around 240 m and 396 m, respectively. We focused on quantifying the amount of  
271 gas emitted in these areas (Sec. 4.3 and 4.4). In addition, we found numerous gas  
272 emissions on the shelf at water depths of about 80 to 90 m and particularly from a  
273 ~50 m high ridge (Area 1) that is part of the forlandet morain complex (Landvik  
274 et al., 2005). Gas bubble emissions occurred in clusters on the ridge and even  
275 more flares were recognized close to the rim of the plateau on top of the ridge.

276 In addition to gas emissions in the three main areas (Areas 1-3), flares were found  
277 widespread at the shelf. Those flares occurred more dispersed compared to the  
278 aggregations at the forlandet morain complex and their relative intensity was  
279 generally weak compared to those recorded in Areas 1, 2, and 3. Flares  
280 preferentially occurred on topographic highs such as shelf break-parallel ridges  
281 that we interpret as recessional moraines. It should be noted, however, that the  
282 distribution of the flares as shown in Figure 2 is biased by the survey line spacing.  
283 Dense line spacing increases the chance to hit a bubble emission, therefore, the  
284 track line of the ship is plotted in Figure 2 as well. Another topographic feature on  
285 the shelf with a considerable number of gas emissions is the transverse ridge at

286 the northern border of the Isfjord cross-shelf trough. More survey lines would be  
287 needed to unravel if this feature might also be a significant source region for gas  
288 emissions.

289 We found no evidence for gas bubble emission in Area 4 (Fig. 1) connected to  
290 pockmarks, which are rounded to elongated depressions at the seafloor at depths  
291 between about 800 and 1200 m. Sixteen pockmarks were crossed during our  
292 hydroacoustic surveys but flares have not been detected in the EK 60 records.  
293 While pockmarks are generally considered as traces of cold fluid seepage, we  
294 conclude that gas bubble emission was not active at the time of investigation.

295 The ~396 m depth contour is the relevant depth, where flares would be expected  
296 to occur, if one or both of the hypotheses of global-warming induced hydrate  
297 dissolution or a seasonal GHSZ are correct. Therefore, we expanded our survey  
298 along this depth for about 80 km to the north (Area 5). However, during this  
299 survey we found no evidence for bubble emissions neither in the EM 710 nor in  
300 the EK 60 records suggesting that the 396-m flares were restricted to Area 3 west  
301 of Prins Karls Forland.

302

#### 303 **4.2 ROV-based observations and vent gas composition**

304 In total we conducted nine remotely operated vehicle (ROV) dives in Areas 1, 2,  
305 and 3 (Table 1). The seafloor at Area 1 (80 to 90 m water depth), that is located at  
306 the main ridge of the forlandet moraine complex, was composed of cobble to  
307 boulder-sized rocks (Fig. 3A) that we interpret as glacial till. Fine grained  
308 sediment filled the space between rocks. Bivalve shells, living sea urchins and  
309 other hardground biota were observed. Bubble emission sites in Area 1 were  
310 patchily distributed. Bubbles rose through rocks or fine grained sediments, with,  
311 in the latter case, whitish microbial mats associated.

312 In Area 2 (240 to 245 m water depth) the proportion of soft sediment was higher  
313 compared to Area 1. However, similar to Area 1 cobble to bolder-sized rocks of  
314 glacial origin occurred. In addition, rocks resembling methane-related  
315 authigenic carbonates were found associated to bubble streams. Bubbles were  
316 released from cm-sized fractures. In places crusts were fractured exposing cavities  
317 below the crust (Fig. 3B). At some sites bubbles accumulated below crusts

318 leading to a periodic release of bursts of bubbles alternating with times of  
319 quiescence. Microbial mats were observed on soft sediments and around bubble  
320 emissions on hard ground.

321 In Area 3 ('396-m flares'), ROV dives were carried out at three locations. In  
322 general, the proportion of soft sediments again was higher compared to that at the  
323 shallower sites. As found in Area 2, crusts resembling methane-related authigenic  
324 carbonates were present. Microbial mats occurred around bubble emission sites  
325 on rocks and on soft sediments. Pogonophoran tubeworms (Siboglinidae) covered  
326 by microbial mats were observed (Fig. 3F). Swarms of demersal fish were  
327 encountered.

328 Analysis of the composition of gas bubbles sampled with the GBS at six bubble  
329 streams in all three areas showed that the gas from Areas 2 and 3 is generally  
330 dominated by methane (99.70 to 99.99% ( $\Sigma(C_1-C_3, CO_2)$ ); Table 2). Only the  
331 single gas sample from Area 1 (90 m) contained a noticeable fraction of  $CO_2$   
332 (~1%). The  $C_1/C_2$  ratio of all samples ranged between 7800 and 15000.

333

### 334 **4.3 Quantification of gas fluxes in Area 2 (240 - 245 m)**

335 In order to conduct an order-of-magnitude estimate of the flux of gas emitted in  
336 Area 2, we followed a simple approach: at first, we quantitatively mapped flares  
337 using the water column data acquired with EM 710. During ROV-dives we found  
338 out that bubble streams occurred in cluster. While one bubble stream may be  
339 enough to cause a flare in several instances more than one stream was  
340 encountered in most cases. We therefore studied several clusters and estimated the  
341 number of bubble streams per cluster. Finally, we estimated the flux of methane  
342 emitted per bubble stream. We then estimated the flux of methane for the entire  
343 area by conducting minimum and maximum estimations that encompass a wide  
344 range of uncertainty.

345 In order to quantitatively map flares in Area 2, we used the water column data  
346 recorded by EM 710 as illustrated in Figure 4. The EM 710 survey was designed  
347 in such a manner that almost complete coverage of the area (gray-shading in  
348 Figure 5) was achieved while significant overlap could be avoided. In total, 512  
349 flares originating from the seafloor in about 240 to 245 m water column were

350 picked from the EM 710 water column data. Most flares concentrated along  
351 lineaments trending parallel to the shelf break. The shelf in this area is flat  
352 without discernable morphology based on the swath bathymetry.

353 Flare intensities varied, but due to noisy EM 710 data classification of flare  
354 intensities (weak vs. strong) could not be achieved, this was left to the EK60 data.  
355 Two ROV dives were conducted in Area 2 (Fig. 6) at sites where weak and strong  
356 flares occurred close to each other (Table 1). For practical reasons, we termed a  
357 site where we found one or more gas emissions within a small area a 'cluster'.  
358 The appearance of cluster C6 in the sonar record is shown in Figure 7. Within a  
359 distance of less than  $\sim 3$  m, we observed 5 bubble streams (S1-S5). We assumed  
360 that all these bubble streams contributed to a flare imaged with EM 710 because  
361 the distance between the streams (max. 3 m) was smaller than the footprint size of  
362 the EM 710 (about 5 m; 1.7% of water depth). In total, we found six clusters  
363 composed of 1 to 15 bubble streams (average  $\sim 6$ ) in Area 2 (Table 3).

364 At 15 individual bubble emission sites (at 5 different clusters) we either calculated  
365 the gas volume flux by interpreting ROV-based videos (visual quantification) or  
366 measured it by placing an inverted funnel (bubble catcher) over the streams (Figs.  
367 3 C and D). Application of both methods at two emission sites showed that the  
368 differences were less than 25% (Table 3). On average,  $15.2 \text{ ml min}^{-1}$  of gas (std.  
369 dev. =  $7.5 \text{ ml min}^{-1}$ ,  $n = 15$ ) were emitted from an emission site. Assuming that  
370 the bubbles consisted of pure methane these rates correspond to methane flux  
371 rates of  $17 \pm 8 \text{ mmol min}^{-1}$ .

372 Based on the flux rates mentioned above, we estimated the flux of methane as gas  
373 bubbles from the seafloor for the entire Area 2. Multiplying the number of 512  
374 known flares existing in Area 2 with average numbers of 6 individual bubble  
375 streams per cluster (Table 3), and average methane flux rates at each bubble  
376 stream ( $17 \text{ mmol min}^{-1}$ ), and assuming that the gas is pure methane,  $52 \text{ mol CH}_4$   
377  $\text{min}^{-1}$  are emitted in Area 2.

378 We further estimated minimum and maximum flux rates by considering the  
379 uncertainties inherent to the approach. An uncertainty of more than one order of  
380 magnitude is introduced by the number of bubble streams feeding a flare as it  
381 varied between 1 and 15 (Table 3). The variability of the flux of a bubble stream

382 ( $17 \pm 8 \text{ mmol min}^{-1}$ ) is comparably small (less than factor 2). Furthermore, we  
383 regard other potential sources of errors not detailed here as comparably negligible.  
384 Calculated minimum and maximum fluxes, which solely considered that between  
385 1 and 15 bubble streams were found to feed a flare, resulted in flux rates ranging  
386 between  $9$  and  $130 \text{ mol min}^{-1}$ , respectively. Assuming a constant flux over time,  
387 the above mentioned values translate to  $27 \times 10^6$  (min:  $5 \times 10^6$ , max:  $68 \times 10^6$ )  
388  $\text{mol CH}_4 \text{ yr}^{-1}$ .

389

#### 390 **4.4 Quantification of gas fluxes in Area 3 ('396-m flares')**

391 We quantitatively looked for gas emissions with the EM 710 in Area 3 at the  
392 upper continental slope (Fig. 8). The distribution of flares was similar to early  
393 observations of Westbrook et al. (2009) and confirmed that the majority of flares  
394 are located at an interval between 360 and 415 m water depths.

395 Preliminary results during our cruise revealed that flares were difficult to pick in  
396 the EM 710 data as they were not stable over time and due to the fact that the  
397 location of flares at the seafloor varied. Therefore, we used a statistical approach  
398 as we were mainly interested in the question of how many flares occur in Area 3  
399 at any given time. For this approach we used four equally spaced hydroacoustic  
400 profiles running across the area where most flares group together. By plotting all  
401 flare positions picked from the EM 710 record (Fig. 8), we identified that more  
402 than 90% of the flares detected in Area 3 occurred in a restricted NW-SE trending  
403 'seep area' (Fig. 8). We used the data obtained during the four transects crossing  
404 this 'seep area' to determine the number of flares during each crossing (Fig. 9).  
405 Because each crossing covered only part of the 'seep area' we calculated the total  
406 number of flares by assuming that the flares were regularly distributed.  
407 Subsequently, we counted the number of flares within the observed area, which is  
408 the seep area within the footprint of the EM 710 (e.g. the red rectangle in Fig. 9A)  
409 and extrapolate that number to the entire seep area (Table 4). The resulting  
410 average number of flares within the 'seep area' was 452. The observed range  
411 (min. = 384, Fig. 9D; max. = 524, Fig. 9B) gave an indication of the uncertainty  
412 inherent to the methodology used and the variability of gas emissions.

413 The temporal variability of bubble emissions was confirmed during ROV dives.  
414 We found that individual bubble streams were transient with bubbles being  
415 emitted for seconds or tens of seconds followed by minutes of inactivity. In  
416 addition, the sites of emission changed spatially within a few decimeters. We  
417 estimated the number of bubble streams occurring in cluster by observing the area  
418 using the horizontally looking sonar for several minutes per site and counted the  
419 number of streams that became visible during the observation time. The numbers  
420 given in Table 5 reflect maximum values: at a given moment bubbles were  
421 emitted only from some sites, i.e. only a fraction of the total number of emission  
422 sites was active. The quantified volume flux at several bubble streams resulted in  
423  $20.9 \text{ ml min}^{-1}$  on average (Table 5). The high variability is reflected in a large  
424 standard deviation of  $15.9 \text{ ml min}^{-1}$  ( $n = 8$ ). The values correspond to a mass flux  
425 of  $18.3 \pm 9.1 \text{ mmol min}^{-1}$  assuming pure methane.

426 The total seafloor flux of methane in Area 3 was calculated based on the  
427 following numbers: Considering average numbers of flares ( $n = 452$ ) and of  
428 bubble streams per cluster ( $n = 6$ ) and an average  $\text{CH}_4$  mass flux ( $18.3 \text{ mmol min}^{-1}$ ),  
429 about 50 moles of methane per minute are emitted in Area 3. Because the  
430 uncertainty inherent to this approach is expectedly large, we conducted  
431 estimations of the minimum and maximum flux. If we consider that only 384  
432 flares occur in Area 3 (Table 4) and assume that each flare may be sourced by a  
433 single bubble stream with an average  $\text{CH}_4$  mass flux only, this results in a seafloor  
434 methane flux of  $7 \text{ mol min}^{-1}$  in Area 3. Calculation of the maximal flux  
435 considering the maximum numbers of flares ( $n = 523$ ) and of bubble streams  
436 found in a cluster ( $n = 10$ ) and average mass fluxes, resulted in  $96 \text{ mol CH}_4 \text{ min}^{-1}$   
437 in Area 3. These values correspond to fluxes of  $26 \times 10^6$  (min.  $4 \times 10^6$ , max.  $50 \times$   
438  $10^6$ )  $\text{mol CH}_4 \text{ yr}^{-1}$ .

439

## 440 **5 Discussion**

### 441 **5.1 Sources of methane**

442 Traditionally, light hydrocarbons of microbial and thermogenic origin are  
443 distinguished by the relation of their molecular composition and the methane  
444 stable carbon isotope ratio (e.g. Whiticar, 1990). The molecular composition of

445 gas in bubbles collected with the GBS several centimeters above the seafloor in  
446 Areas 2 (240 to 245 mbsl) and Area 3 ('396-m flares') indicate a predominantly  
447 microbial origin of the vent gas ( $C_1/C_2$  ca. 9,700 to 15,200; Fig. 10). However,  
448 less negative  $\delta^{13}C\text{-CH}_4$  ratios ( $-53.8$  to  $-57.4\text{‰}$  V-PDB) than expected from the  
449 molecular composition for typical microbial methane point to some admixture of  
450 methane enriched in  $^{13}C$ . A possible explanation for this observation might be that  
451 part of the methane has undergone oxidation within the sediments, which would  
452 result in  $^{13}C$ -enrichment of the residual methane.

453 Our finding of gas with an average  $\delta^{13}C$  ratio of  $-55.7\text{‰}$  in Areas 2 and 3  
454 complements well results from water column studies in Area 2 carried out by  
455 Gentz et al. (2014). Using correlations between concentration and stable carbon  
456 isotopic compositions of methane in the water column the authors inferred the C-  
457 isotope signature of methane emitted from the seafloor (about  $-60\text{‰}$ ). A similar  
458  $\delta^{13}C$  ratio ( $-54.6 \pm 1.7\text{‰}$ ) was reported by Fisher et al. (2011) for methane in  
459 hydrates recovered from an area termed 'Plume field' (890 m water depth), which  
460 is identical to our Area 4. In summary, the source of methane at the upper  
461 continental slope and outer shelf (Areas 2, 3, and 4) appear to be similar based on  
462 its geochemical signature and largely microbial in origin.

463 Gas emitted as bubbles at the shelf in Area 1 ( $\sim 90$  m water depth) differs from  
464 that sampled in Areas 2 and 3 in its molecular composition ( $C_1/C_2$  ca. 7,850) and  
465  $\delta^{13}C\text{-CH}_4$  ratio ( $-43.5\text{‰}$  V-PDB) (Fig. 10). This difference is significant, but only  
466 a single gas sample could be obtained from Area 1 during our research cruise.  
467 Nevertheless, this finding generally agrees with the water column study by Damm  
468 et al. (2005) carried out on a much larger scale along the entire SW continental  
469 margin of Svalbard. The authors postulated widespread methane seepage along  
470 the shelf with respect to methane enrichments at several stations. In addition, the  
471 authors observed a topography-dependent methane isotope signature with  $-30\text{‰}$   
472 at the tops and  $-49\text{‰}$  in troughs. Damm et al. (2005) conclude that the  
473 geochemical signature of methane is influenced due to its slow seepage through  
474 the sediments leading to 'inter-granular seepages or micro-seepages'. Our results  
475 clearly show that methane emission at the shelf is not limited to micro-seepage,  
476 but also occurs as vigorous bubble emission as observed at the main ridge of the  
477 forlandet moraine complex.

478 Unfortunately, our sparse results on the gas composition and methane isotope  
479 signature at the forlandet moraine complex do not allow any final assessment of  
480 the source of methane (Fig. 10) because migration, oxidation, and in situ  
481 generation of gas might have overprinted the original signature. Additional gas  
482 samples (e.g. from the deeper subsurface) are needed to ultimately clarify this  
483 aspect.

484

## 485 **5.2 Distribution of gas emissions at the seafloor**

486 The results of our extensive hydroacoustic survey (single beam and swath  
487 mapping) provide valuable insight into the system of gas emission at the  
488 continental margin west of Svalbard. We have covered large areas searching for  
489 flares with hydroacoustic techniques (Fig. 1), but evidence for gas emissions was  
490 restricted to the region west of Prins Karls Forland. This region is apparently  
491 prone to fluid flow as suggested by gas emissions occurring all over the shelf and  
492 upper slope. Gas emissions exclusively occur in this inter-fan region bordered by  
493 the Kongsfjord cross-shelf trough to the north and the Isfjord cross-shelf trough to  
494 the south.

495 The swath bathymetry acquired during our cruise significantly extends published  
496 maps (Landvik et al., 2005; Ottesen et al., 2007) and shows a series of along-  
497 shelf, parallel ridges between the shelf break and the forlandet moraine complex  
498 (Fig. 2). We interpret these ridges as surface expressions of prograding foresets,  
499 which are sediments deposited at the seaward termination of ice sheets during  
500 phases of progression and regression. Because seismic data acquired in the region  
501 comprising Areas 2 and 3 show prograding glacial sequences at the outer shelf  
502 (Rajan et al., 2012; Sarkar et al., 2012), it can be expected that these also occur  
503 further to the south. Gas emissions occur all over the shelf with a peculiar  
504 clustering at the forlandet moraine complex. In contrast, the distribution of gas  
505 emissions at the shelf distant to the forlandet moraine complex does not follow  
506 any discernable pattern; however, there might be a weak tendency that flares  
507 preferentially occur at topographic highs but not in depressions.

508 Numerous flares concentrated at the forlandet moraine complex at water depth of  
509 about 80 to 90 m (Fig. 2). The detailed hydroacoustic surveys conducted during

510 our cruise revealed that almost all flares originated from the top of the moraine,  
511 which suggests that the methane source might be located within the  
512 morphological ridge itself. However, as we lack data on the sub-seafloor  
513 structure, this remains speculative. Potential capacious methane reservoirs at  
514 Arctic continental shelves are methane-loaded sediments below permafrost (e.g.  
515 Rachold et al., 2007). Transgression of the ocean following the last glacial stage  
516 has led to submergence and subsequent dissolution of permafrost in the sediments  
517 induced by bottom-water temperatures  $>0^{\circ}\text{C}$ . In case the permafrost seal is  
518 broken, methane can escape the reservoir and may be emitted as bubbles from the  
519 seafloor, a process recently observed on large scales at the East Siberian Shelf  
520 (Shakova et al., 2010). Still ongoing permafrost melting may, thus, be an  
521 explanation for the concentrated gas emissions observed at the forlandet moraine  
522 complex. In case this holds true, a microbial origin of the expelled gas would be  
523 expected. Unfortunately, the geochemical properties of the gas sample collected  
524 in Area 1 do not allow for unambiguous source assignments. Additional sub-  
525 surface gas samples are needed to unravel the gas source at the forlandet moraine  
526 complex.

527 Flares in Areas 2 and 3 are potentially sourced by dissociating gas hydrates  
528 (Westbrook et al., 2009; Berndt et al., 2014). Bubbles in Area 2 are emitted at  
529 shallow depth of about 240 to 245 m above the GHSZ. Seismic studies, however,  
530 have shown that the flares may also be sourced by dissociating hydrates (Rajan et  
531 al., 2012; Sarkar et al., 2012). A slump at the upper slope and prograding forsets  
532 led to the landward deviation of upward migrating fluids, such that the gas is  
533 emitted along lineaments at the outer shelf (Fig. 5).

534 Flares in Area 3 are linearly orientated along a band at the upper continental slope  
535 at water depth above  $\sim 396$  m (Fig. 8). Using the swath echosounder, we  
536 systematically mapped the upper slope in order to quantitatively record the  
537 occurrence of flares in Area 3. In accordance with earlier observations we found  
538 that the majority of gas emissions occurred along a narrow band (gray shaded  
539 'seep area' in Fig. 8) with some additional flares located above and below that  
540 area, a pattern that was attributed to small-scale lithological heterogeneity before  
541 (Sarkar et al., 2012).

542 While our results do not allow to conclude whether methane emissions in Area 3  
543 are fed by dissociating gas hydrates, we are able to refine the depth-dependent  
544 flare distribution already proposed before Westbrook et al. (2009) with our data.  
545 The abundance of flares versus depth in Area 3 is shown in Figure 11. Because  
546 the depth-related abundance of flares resembles a Gaussian distribution a generic  
547 link between depth and gas emission is intuitive.

548 Because most flares occurred between about 360 and 415 m water depth it is  
549 tempting to calculate the sediment temperature increase which would be required  
550 to induce hydrate dissociation. For this, we calculated the gas hydrate phase  
551 boundary using the composition of gas sample GeoB 16833-2 collected with the  
552 GBS (Fig. 11). The resulting increase in sediment temperature of 1.2 °C is in  
553 agreement with both hypothesis proposed to explain the narrow zone of flare  
554 origins at the seafloor: a 1°C temperature increase during the last 30 years  
555 (Westbrook et al., 2009) and a seasonal fluctuation of 1–2 °C as measured with  
556 the MASOX lander (Berndt et al., 2014).

557 Based on the seafloor flare distribution determined in this study, we conclude that  
558 if gas hydrate dissolution is a cause for seafloor gas emissions, this process was  
559 spatially limited to one segment at the continental margin (west of Prins Karls  
560 Forland) during the time of our investigation. Furthermore, the presence of  
561 numerous additional flares at the shelf suggests that this particular region west of  
562 Prins Karls Forland is prone to hydrocarbon seepage and that gas seafloor  
563 emission unaffected by gas hydrate dissociation is common in the region.

564

### 565 **5.3 Quantification of gas bubble emissions**

566 Combining hydroacoustic data with ROV-based observations, we quantified the  
567 flux of methane as gas bubbles from the seafloor to the water column. This  
568 approach is advantageous because it is relatively simple and straight forward  
569 providing order-of-magnitude estimations for gas bubble fluxes. Similar  
570 methodologies were recently applied in other settings characterized by gas bubble  
571 emissions (Römer et al., 2012a; Römer et al., 2012b; Römer et al., 2014; Sahling  
572 et al., 2009).

573 Here, we discuss two major sources of uncertainty in our flux calculations that we  
574 regard as most important. Our estimation is a snapshot in time, taken at a few  
575 days in Aug/Sept 2012. This is especially important in light of the recently posed  
576 hypothesis (Berndt et al., 2014) that a temperature-induced annual build-up and  
577 break-down of hydrates would lead to an annual cycle in the gas emissions. Our  
578 results show that the gas emissions were persistent for hours (ROV-observations)  
579 or even days (repeated hydroacoustic observations, Fig. 9; Tab. 4). In addition,  
580 gas emissions were encountered each year since their discovery in 2008  
581 (Westbrook et al., 2009): 2009 (Fisher et al., 2011; Rajan et al., 2012), 2010  
582 (Gentz et al., 2014), 2011 (Wright, 2012), 2012 (Berndt et al., 2014; this study).  
583 All investigations of gas emissions in that region so far were carried out in the  
584 summer period, and, therefore, it is uncertain whether the gas emissions undergo  
585 annual periodicity. In order to test the hypothesis by Berndt et al. (2014), a  
586 research campaign in spring, when bottom water temperatures are minimal and  
587 the thickness of the GHSZ should peak (and thus bubble emission may be  
588 minimal), would be useful. In this study, we state gas fluxes per year for  
589 comparative purposes (see below) although the temporal variability of gas  
590 emissions is unknown.

591 Our quantification approach revealed a source of uncertainty that waits for a  
592 technical solution, i.e. an answer to the question how many individual streams of  
593 bubble contribute to one flare as imaged by ship-mounted multibeam  
594 echosounder. By use of the ROV-mounted horizontally-looking sonar we found  
595 that a single bubble stream is enough to cause a flare but that sometimes up to 15  
596 bubble streams contribute to one flare (Table 3). While the bubble flux of a single  
597 bubble stream can appropriately be determined by using a ROV (visually or by  
598 capturing the bubbles), and the numbers of flares can be systematically mapped  
599 using multibeam, the uncertainty in the bubble stream-to-flare ratio introduces a  
600 factor of  $>10$ . In this study, we employed the ROV-mounted sonar for this  
601 purpose but encountered several shortages, i.e. the difficulty to keep the ROV  
602 stationary at strong bottom-water currents and the need of very long scanning  
603 times consuming a lot of highly valuable ROV operation time. A towed sonar  
604 system or a sonar on a bottom-mounted lander system would be desirable  
605 technical innovations.

606 The bubble flux of methane in Areas 2 (5 to  $68 * 10^6$  mol yr<sup>-1</sup>) and 3 (4 to  $50 * 10^6$  mol yr<sup>-1</sup>) estimated in this study, is similar to the range of fluxes (0.23 to  $87 * 10^6$  mol yr<sup>-1</sup>) in other bubble emission settings (Table 6). Because bubble fluxes in all these settings are in the same order of magnitude gives confidence that our approach used for estimating the flux in this study is reliable.

611 Our estimation of the bubble flux contributes to the ongoing discussion about the amount of gas hydrate in the upper continental slope west of Svalbard that is susceptible for temperature changes. We base the following discussion on the assumption that most of the methane is released as gas bubbles from the seafloor, when hydrates within the seafloor are dissociating. We neglect the amount of methane that is consumed by oxidation within the seafloor or that is emitted dissolved in the aqueous phase, as we have no control on these processes. In order to compare flux rates determined in this study with those given in the literature for hydrate dissociation, we converted published rates into the annual methane flux per meter margin segment (mol m<sup>-1</sup> yr<sup>-1</sup>). Our systematic flare mapping revealed that the gas emission-influenced margin segment has a length of ~14 km (Areas 2 and 3, Fig. 5 and 8), which is short compared to those (30 and 25 km, respectively) investigated in other related studies (Westbrook et al., 2009; Marín-Moreno et al., 2013).

625 Overall, the bubble flux estimated in this study is lower than the amount of methane released from dissociating hydrates reported earlier (Table 7). However, the published rates span three orders of magnitude with minimum rates being consistent with our estimates. Westbrook et al. (2009) initially estimated methane release from dissociating hydrates at about  $56 * 10^3$  mol m<sup>-1</sup> yr<sup>-1</sup>. Based on 2D modeling (Reagan et al., 2011) scaled this value down to  $8.8 * 10^3$  mol m<sup>-1</sup> yr<sup>-1</sup>, which is about the same order of magnitude as our bubble-flux estimate (0.6 to  $8.4 * 10^3$  mol m<sup>-1</sup> yr<sup>-1</sup>). These fluxes are based on an increase in bottom-water temperature of about 1°C during the past three decades considering progressive hydrate dissolution at present. If the gas emission in Areas 2 and 3 are sourced by temperature-induced multi-year hydrate dissolution, the model by Reagan et al. (2011) appear to be most applicable.

637 The impact of future bottom-water warming on hydrates in sediments of the upper continental slope west off Svalbard was modeled by Marín-Moreno et al. (2013)

639 using climate models and scenarios representing low and high greenhouse  
640 emissions (i.e. representative concentration pathways 2.6 and 8.5, respectively).  
641 During the upcoming 100 years, the hydrate dissolution rate is forecasted at 6.9 to  
642  $20.6 * 10^3 \text{ mol m}^{-1} \text{ yr}^{-1}$  with acceleration to  $13.2$  to  $72.3 * 10^3 \text{ mol m}^{-1} \text{ yr}^{-1}$  within  
643 the next 300 years. These rates are, again, higher compared to those determined  
644 for present bubble emissions in this study. The predictions by Marín-Moreno et al.  
645 (2013) call for monitoring of the hydrate deposits west of Svalbard in the future.

646 According to Berndt et al. (2014) release of methane from the dynamic hydrate  
647 reservoir amounts to  $561$  to  $935 * 10^3 \text{ mol m}^{-1} \text{ yr}^{-1}$ , which is two orders of  
648 magnitude higher than the bubble flux that we estimate. The discrepancy between  
649 these values warrants further investigation. A first approach could be, to test  
650 whether bubble emission intensities actually vary during the year.

651 Because methane is a potent greenhouse gas, the fate of methane emitted from the  
652 seafloor is of relevance. Gentz et al. (2014) showed for the well-stratified water  
653 column in Area 2 during the summer that the majority of methane is diffusing  
654 from bubbles into the water column below the pycnocline and leads to relative  
655 enrichments in the concentrations of dissolved methane in the lower water body.  
656 However, as the lower water body is isolated from the upper water layer by the  
657 density difference (the pycnocline), methane dissolved in lowermost water masses  
658 does not reach the atmosphere. Therefore, most of the methane emitted from the  
659 seafloor is either oxidized, or transported in the water mass and further diluted, or  
660 reaches the sea surface, where it could escape into the atmosphere. Complete  
661 methane removal by oxidation occurs within about 50 to 100 days (Gentz et al.,  
662 2014). Therefore, the fate of methane depends on the timeframe and fate of the  
663 water mass. The situation is different in autumn, when storms and low  
664 temperatures break down the water column stratification and induce vertical  
665 mixing. Although not studied so far, it might be expected that bubble-forming  
666 methane gets dissolved in the water and transported through the water-air  
667 interface into the atmosphere, contributing to the atmospheric methane inventory.

668

669 **6 Conclusion**

670 At the upper slope (Area 3) and outer continental shelf (Area 2) methane of  
671 microbial origin is emitted at the seafloor. Based on our data, we cannot  
672 contribute to the question if gas hydrate dissolution is the cause for the observed  
673 bubble emissions and, further, if a multi-year warming trend (1°C in 30 years) or  
674 a seasonal temperature cycle is the driver of the hydrate dissolution. But our data  
675 show that if hydrate dissolution in Areas 2 and 3 occurs, it is spatially limited to a  
676 margin segment of about 14 km and does not occur along the ~80 km 396-m  
677 isobath to the north. Our quantification of gas emissions in Areas 2 and 3 reveals  
678 methane fluxes in the same order of magnitude as found at bubble vents in other  
679 geological settings. If hydrate dissociation is involved, our flux estimate may help  
680 to refine models on this temperature-susceptible reservoir and serves as baseline,  
681 in the case that warming leads to intensified gas emissions in future.

682 The gas emissions in Areas 2 and 3 are only one aspect of fluid flow offshore  
683 Svalbard as bubble vents were found all over the shelf and especially prominent at  
684 the forlandet moraine complex (Area 1) reflecting that the area west of Prins  
685 Karls Forland is prone the gas venting. We speculate that decaying permafrost  
686 may allow methane to escape from a deeper reservoir at the forlandet moraine  
687 complex at water depth around 90 m.

688 **Acknowledgement**

689 We thank the captain and crew of R/V Heincke cruise 387. This work was funded  
690 through the DFG-Research Center/Excellence Cluster “The Ocean in the Earth  
691 System” MARUM – Center for Marine Environmental Sciences. M. T. was  
692 supported by SENSEnet, a Marie Curie Initial Training Network funded by the  
693 European Commission 7<sup>th</sup> Framework Program, Contract No. PITN-GA-2009-  
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838 **Tables**

839 Table 1. Stations and instruments deployed during R/V Heincke cruise HE-387.  
 840 Abbreviations: ROV=remotely operated vehicle MARUM-Cherokee; GBS=Gas  
 841 bubble sampler; Marker=seafloor deployed stone with a syntactic floating foam  
 842 bound to it.

Date	Stat. No.	Stat. GeoB	No. Instrument	Latitude	Longitude	Water depth (m)
23 Aug 2012	7	16807	ROV Dive 01	ca. 78°32.9' N	ca. 10°14.2' E	91
23 Aug 2012	7-1	16807-1	Marker 2	78°32.839' N	10°14.247' E	94
23 Aug 2012	7-2	16807-2	GBS 1	78°32.839' N	10°14.252' E	94
23 Aug 2012	7-3	16807-3	GBS 2	78°32.840' N	10°14.247' E	94
24 Aug 2012	12	16812	ROV Dive 02	ca. 78°32.8' N	ca. 10°14.3' E	83
25 Aug 2012	16	16816	ROV Dive 03	ca. 78°32.8' N	ca. 10°14.2' E	94
27 Aug 2012	23	16823	ROV Dive 04	ca. 78°39.2' N	ca. 9°25.8' E	241
27 Aug 2012	23-1	16823-1	Marker 1	78°39.253' N	9°25.760' E	241
27 Aug 2012	23-2	16823-2	GBS 1	78°39.254' N	9°25.755' E	242
27 Aug 2012	23-4	16823-4	Marker 4	78°39.252' N	9°26.044' E	241
27 Aug 2012	23-5	16823-5	GBS 2	78°39.252' N	9°26.041' E	240
28 Aug 2012	26	16826	ROV Dive 05	ca. 78°39.2' N	ca. 9°26.0' E	243
30 Aug 2012	33	16833	ROV Dive 06	ca. 78°37.1' N	ca. 9°24.6' E	382
30 Aug 2012	33-1	16833-1	Marker 5	78°37.220' N	9°24.659' E	381
30 Aug 2012	33-2	16833-2	GBS 1	78°37.218' N	9°24.659' E	382
30 Aug 2012	33-3	16833-3	GBS 2	78°37.210' N	9°24.570' E	384
30 Aug 2012	33-4	16833-4	Marker 3	78°37.209' N	9°24.565' E	384
02 Sept 2012	46	16846	ROV Dive 07	ca. 78°35.4' N	ca. 9°26.5' E	386
03 Sept 2012	48	16848	ROV Dive 08	ca. 78°33.4' N	ca. 9°28.3' E	391
03 Sept 2012	48-1	16848-1	Marker 8	78°33.334' N	9°28.509' E	387
03 Sept 2012	48-2	16848-2	GBS	78°33.326' N	9°28.558' E	387
04 Sept 2012	53	16853	ROV Dive 09	ca. 78°34.5' N	ca. 10°10.2' E	90

843 Table 2. Proportions of low-molecular-weight alkanes and CO<sub>2</sub> [in mol.% of  
 844  $\Sigma(C_1-C_3, CO_2)$ ] in vent gas samples taken with the Gas Bubble Sampler (b.d.l. =  
 845 below detection limit).  
 846

Area	Depth	ROV Dive	GeoB	CH <sub>4</sub> (mol-%)	C <sub>2</sub> H <sub>6</sub> (mol-%)	CO <sub>2</sub> (mol-%)	C <sub>3</sub> H <sub>8</sub> (mol-%)	C <sub>1</sub> /C <sub>2</sub>	$\delta^{13}C-CH_4$ (‰ V-PDB)
Area 1	90 m	01	16807-2	98.977	0.013	1.009	< 0.001	7852	-43.5
Area 2	240 m	04	16823-1	99.689	0.007	0.303	< 0.001	15161	-55.8
Area 2	240 m	04	16823-3	99.730	0.007	0.261	< 0.001	13919	-55.7
Area 3	380 m	06	16833-2	99.991	0.008	b.d.l.	< 0.001	12213	-53.8
Area 3	380 m	06	16833-3	99.858	0.010	0.131	< 0.001	10325	-57.4
Area 3	380 m	08	16848-2	99.703	0.010	0.286	< 0.001	9697	-56.0

847 Table 3. Gas quantities transported by individual gas bubble streams in Area 2  
 848 (240 – 245 mbsl) determined by use of the gas bubble catcher or by interpretation  
 849 of video footage.  
 850

Cluster	No(s) of bubble streams	ROV - Dive	Location	Tools	Stream	Flux ml /min Visual quantification	Flux ml/min Bubble catcher
C1	15	04	78°39.253'N; 9°25.760'E  241 m	Marker 1, GBS	S1	17.0	
					S2	9.9	
					S4	23.0	
C2	12	04	78°39.252'N; 9°26.044'E  241 m	Marker 4, GBS	S1	21.7	
					S2	13.0	
					S4	8.5	
					S5	6.6	
					S6	20.5	
C3	1	05	78°39.216'N; 9°25.834'E  242 m		S1	26.5	27.9
C4	1	05	78°39.216'N; 9°25.786'E  241 m		S1	5.2	4.0
C5	1	05	78°39.228'N; 9°25.735'E  242 m				
C6	5	05	78°39.201'N; 9°25.995'E  241 m		S1		25.0
					S2		19.4
					S3		6.2
					S4		8.2
					S5		17.1

851 Table 4. Estimated number of flares in Area 3 following the approach described in  
 852 the text and illustrated in Fig. 9.

853

Profile	Observed area	Number of flares in observed area	Ratio observed area to 'seep area' (3.72 km <sup>2</sup> ) in %	Estimated total number of flares in 'seep area'
Fig. 9 A	2.35 km <sup>2</sup>	294	63.1	466
Fig. 9 B	2.89 km <sup>2</sup>	407	77.7	524
Fig. 9 C	2.88 km <sup>2</sup>	334	77.4	432
Fig. 9 D	2.38 km <sup>2</sup>	246	64.0	384
Average				451.5

854 Table 5. Gas quantities transported by of individual gas bubble streams in Area 3  
 855 ('396-m flares') determined by use of the gas bubble catcher and by interpretation  
 856 of video footage.

857

Cluster	Number of bubble streams	Dive	Location Depth	Tools	Stream	Flux ml /min Visual quantification	Flux ml/min Bubble catcher
C1	10	06	78°37.220'N; 9°24.659'E, 385 m	Marker 5, GBS 1	S1	9.4	
C3	3	06	78°37.209'N; 9°24.565'E; 385 m	Marker 3, GBS 2	S1	6.7	
C5	8	07	78°35.380'N; 9°26.627'E; 385 m		S1		6.3
					S2		31.0
					S3		37.5
					S4		41.0
C6	8	07	78°35.381'N; 9°26.604'E; 385 m		S1		3.0
					S2		32.0
C7	4	07	78°35.380'N; 9°26.831'E; 386 m				
C8	5	07	78°33.335'N; 9°28.527'E; 385 m	Marker 8			
C9	6	07	78°33.326'N; 9°28.548'E; 385 m				
C10	3	07	78°33.310'N; 9°28.647'E; 385 m				
C11	6	07	78°33.299'N; 9°28.603'E; 389 m				

858 Table 6. Fluxes of bubble-forming methane from the seafloor to the hydrosphere  
 859 in various regions.

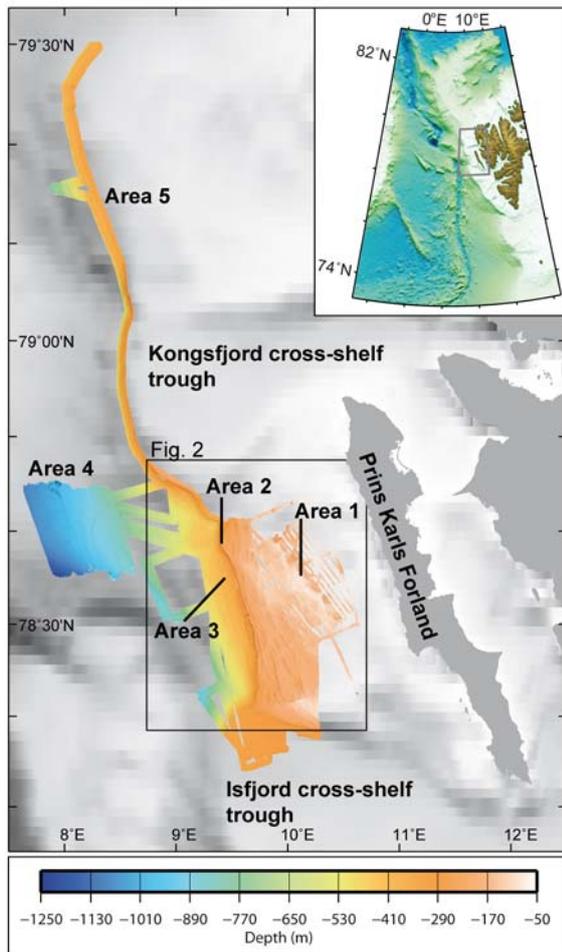
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Methane bubble flux ( $10^6 \text{ mol yr}^{-1}$ )	Water depth (m)	Area	Reference
27 (5 to 68)	240 – 245	Area 2	This study
26 (4 to 50)	380 – 390	Area 3	This study
~19	1250 – 1270	Håkon Mosby Mud Volcano - all three emission sites	Sauter et al., 2006
2 to 87	890	Kerch Flare, Black Sea	Römer et al., 2012a
21.9	600 – 700	Northern summit Hydrate Ridge, offshore Oregon	Torres et al., 2002
1.5	65 – 75	Tommeliten field, North Sea	Schneider von Deimling et al., 2011
40 ( $\pm 32$ )	575 – 2870	Makran continental margin (50 km broad segment)	Römer et al., 2012b
0.23 to 2.3	1690	Carbonate slab, Nile Deep Sea Fan	Römer et al., 2014

861 Table 7. Amount of methane either released as bubbles from the seafloor (this  
 862 study) or susceptible to temperature-induced hydrate dissociation as revealed  
 863 from modeling.  
 864

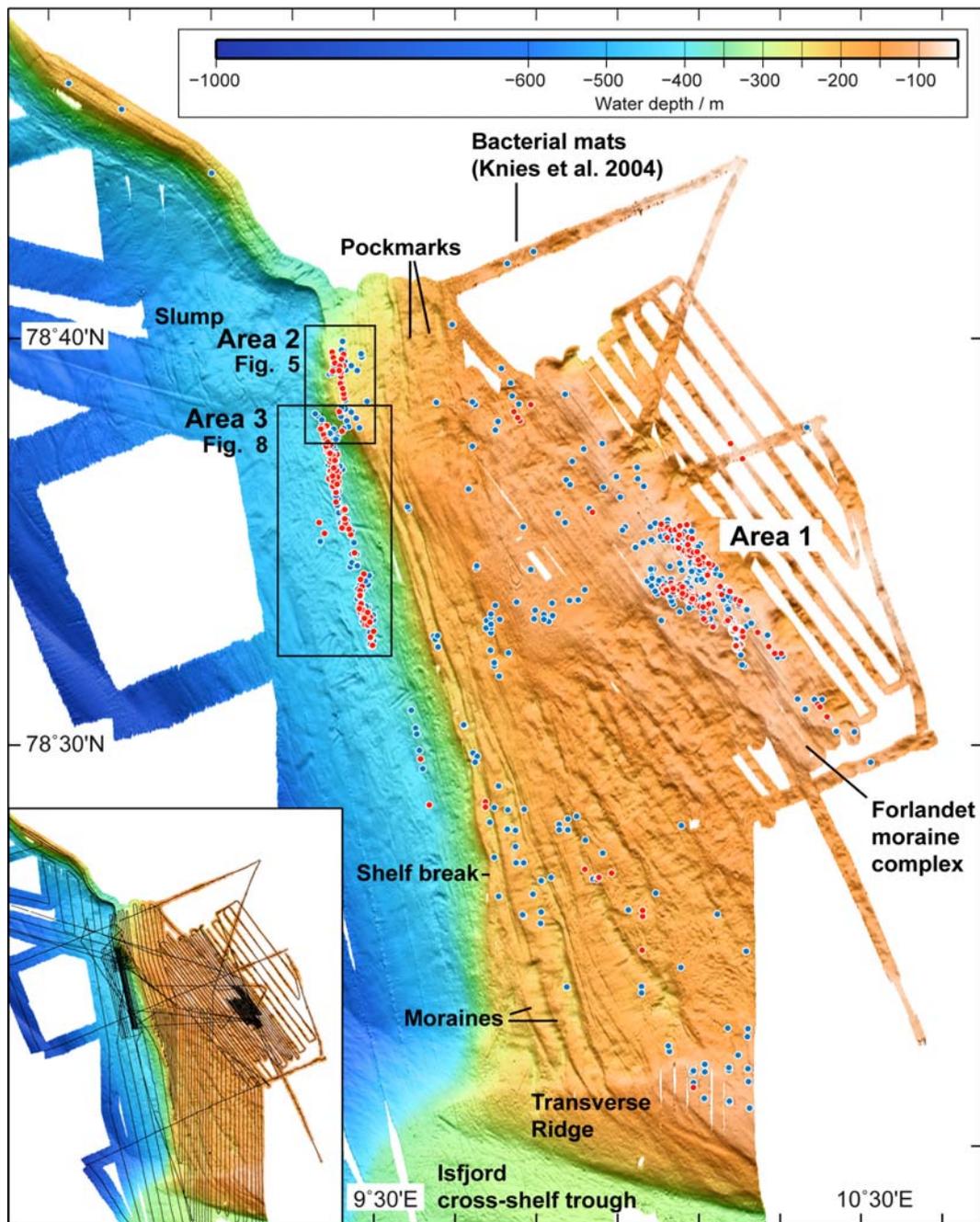
Description	Amount methane ( $10^6$ mol yr <sup>-1</sup> )	Margin width (km)	Amount methane ( $10^3$ mol yr <sup>-1</sup> m <sup>-1</sup> )	Reference
Methane flux as bubbles (Area 2)	27 (5 to 68)	4.5	6.0 (1.1 to 15.1)	This study
Methane flux as bubbles (Area 3)	26 (4 to 50)	11	2.4 (0.4 to 4.5)	This study
Methane flux as bubbles (Area 2 & 3)	53 (9 to 118)	~14	3.8 (0.6 to 8.4)	This study
Progressive dissociation of hydrate	1683	30	56.1	Westbrook et al., 2009
Progressive dissociation of hydrate	264	30	8.8	Reagan et al., 2011
Future (100 years) dissociation of hydrates	171 to 514	25	6.9 to 20.6	Marín-Moreno et al., 2013
Future (300 years) dissociation of hydrates	330 to 1807	25	13.2 to 72.3	Marín-Moreno et al., 2013
Annual hydrate formation and dissociation			561 to 935	Berndt et al., 2014

865 **Figure captions**



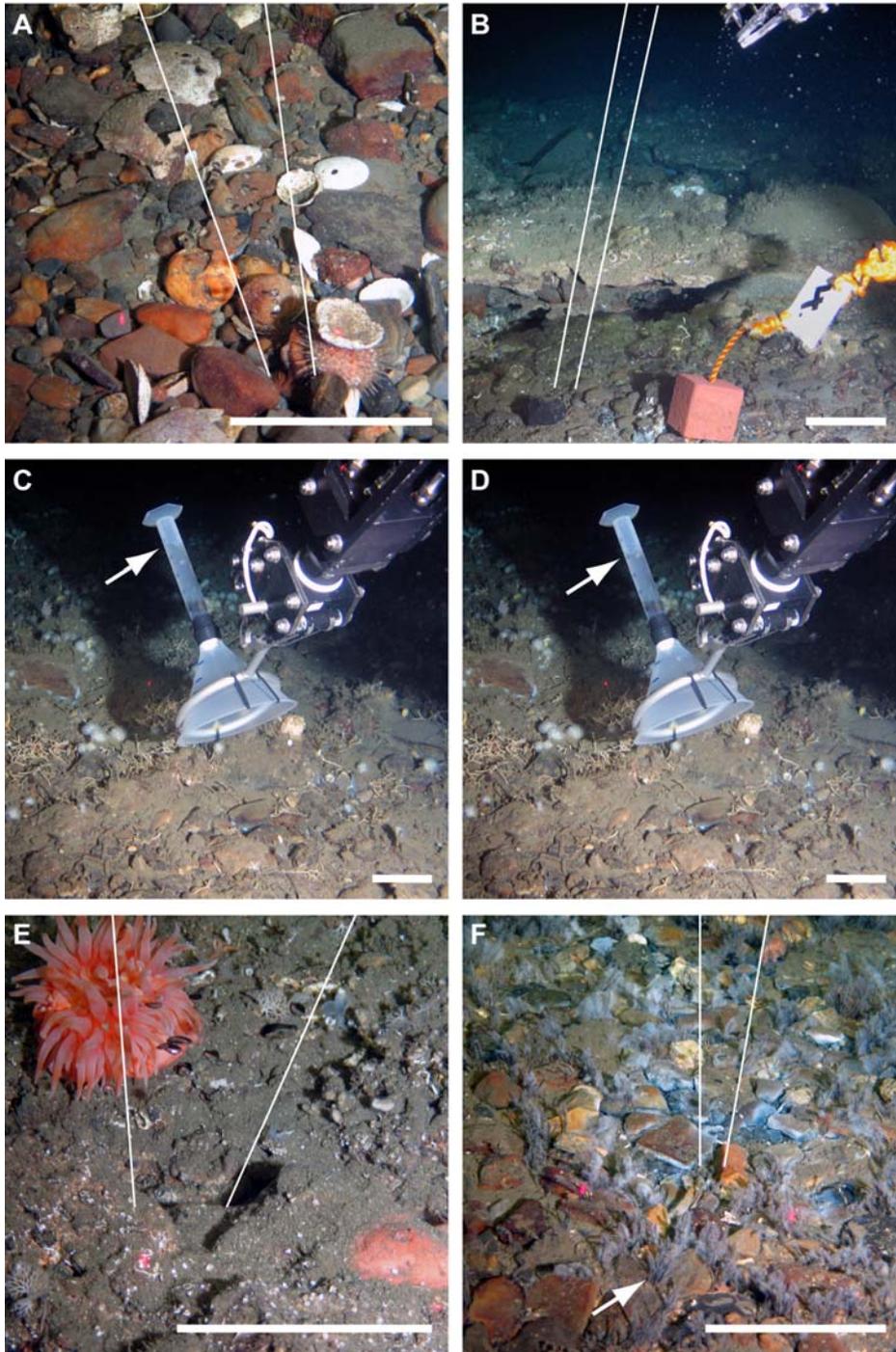
866

867 Figure 1. Multibeam bathymetry obtained during R/V Heincke cruise 387 (colour) plotted  
868 on IBACO bathymetry (Jakobsson et al., 2008) showing the study areas (Areas 1 to 5) at  
869 the continental margin west of Svalbard. Inset shows an overview map with the location  
870 of the main map.



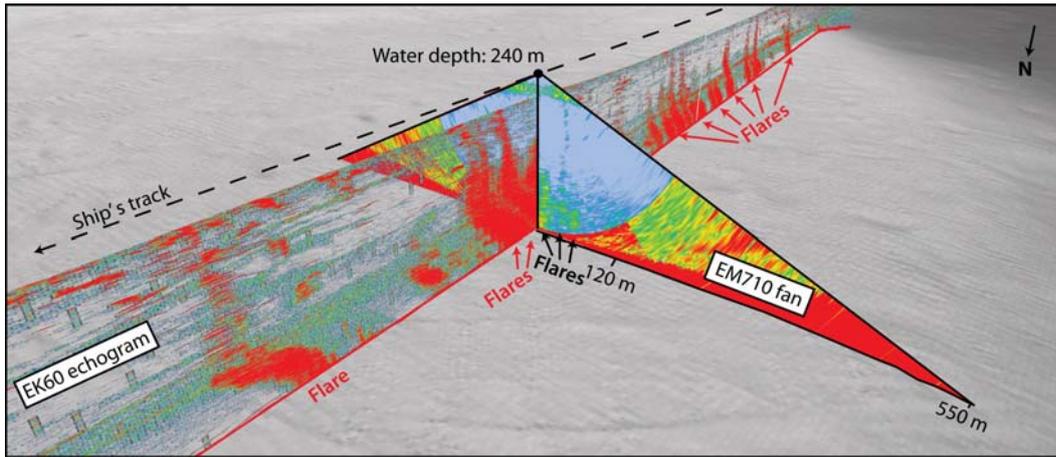
871

872 Figure 2. Main figure: Location of flares (hydroacoustic indications of gas bubble  
 873 emissions) during summer 2012 as picked from EK 60 echosounder records plotted on  
 874 top of multibeam bathymetry. Strong flares (red dots) mainly occur in Areas 1, 2, and 3.  
 875 Weak flares (blue dots) occur widespread at the shelf. Inset: Map of that region showing  
 876 the ship track and bathymetry.



877

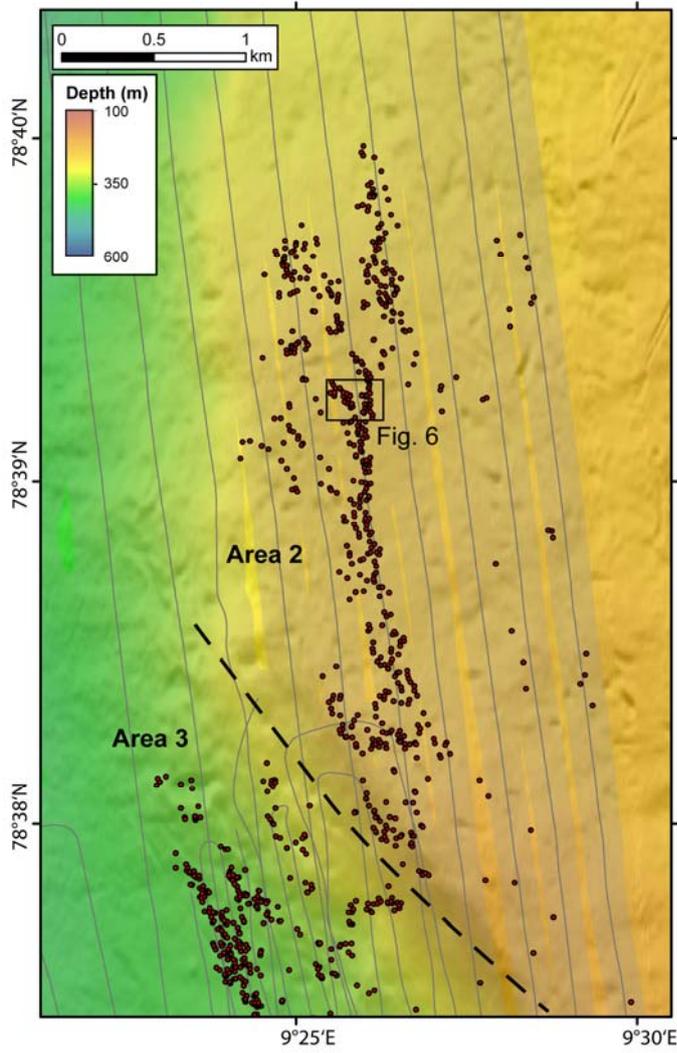
878 Figure 3. Seafloor images taken during dives with ROV at gas emission sites in Area 1  
 879 (A), Area 2 (B-E), and Area 3 (F). Scale bar is 10 cm. Arrows point to objects of interest,  
 880 white lines outline the trace of the rising bubbles. (A) Bubbles escaping from the cobble-  
 881 covered seafloor (Dive 02). (B) Three bubble streams at Marker 4. Crusts resembling  
 882 authigenic carbonates at the seafloor (Dive 04). (C), (D) Images illustrating the use of the  
 883 bubble catcher for measuring the gas bubble volume flux (Dive 05). (E) Bubbles rising in  
 884 front of an anemone (Dive 04). (F) Filamentous (probably sulfur-oxidizing) bacteria and  
 885 pogonophora at a bubble stream (Dive 08). Photos courtesy of MARUM.



886

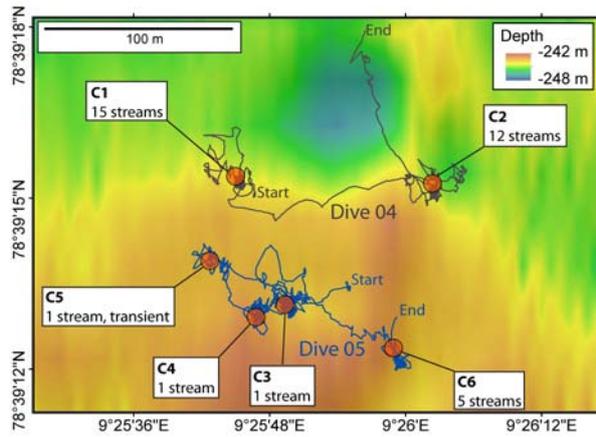
887 Figure 4. Composite figure illustrating the appearance of flares in single beam EK 60  
 888 echosounder and in multibeam EM 710 echosounder. Flares can be traced in the central  
 889 part of the EM 710 fan ( $45^\circ$  to each side), in this example obtained in 240 m water depth,  
 890 the across track width is 120 m to each side. Beyond that limit, the noise is too high to  
 891 reliably map flares.

892



893

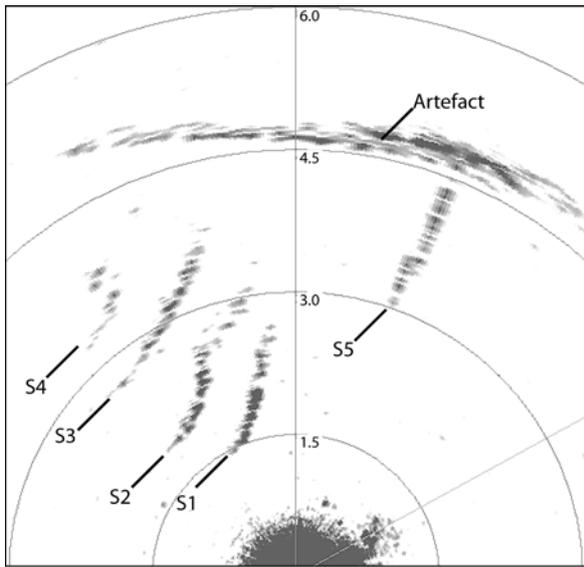
894 Figure 5. Flares (circles) in Area 2 (240 to 245 mbsl) plotted on shaded bathymetry.  
 895 Flares were picked in multibeam water column data; the coverage is shown as grey  
 896 shading around the ship track (lines). ROV dives were performed in an area highlighted  
 897 by the rectangular box (Fig. 6).



898

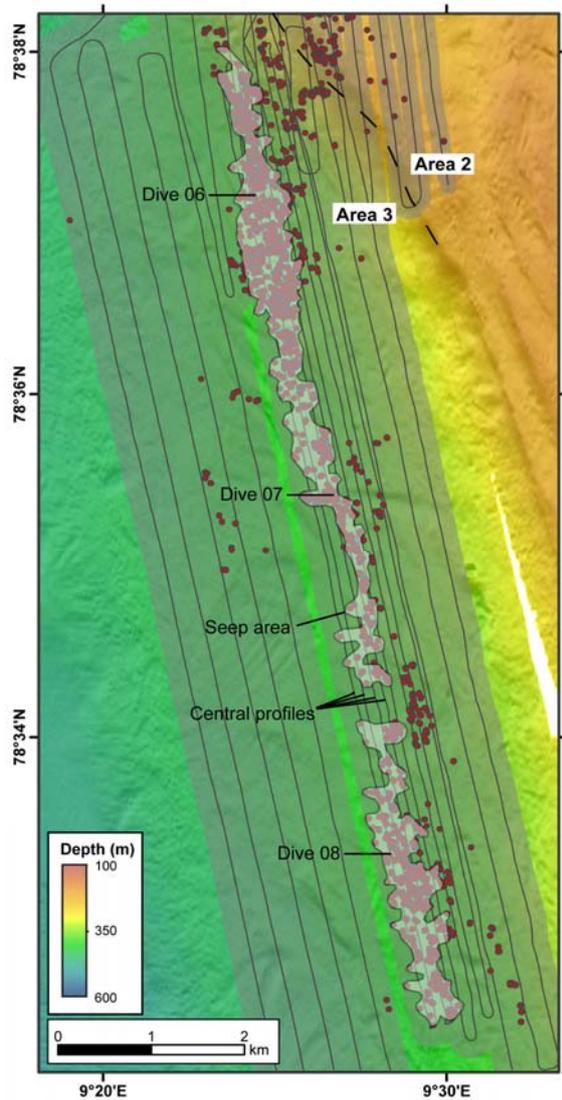
899 Figure 6. Bubble stream clusters (C1 to C6) in Area 2 discovered during ROV dives 04  
 900 and 05. Dive tracks are shown on bathymetry.

901



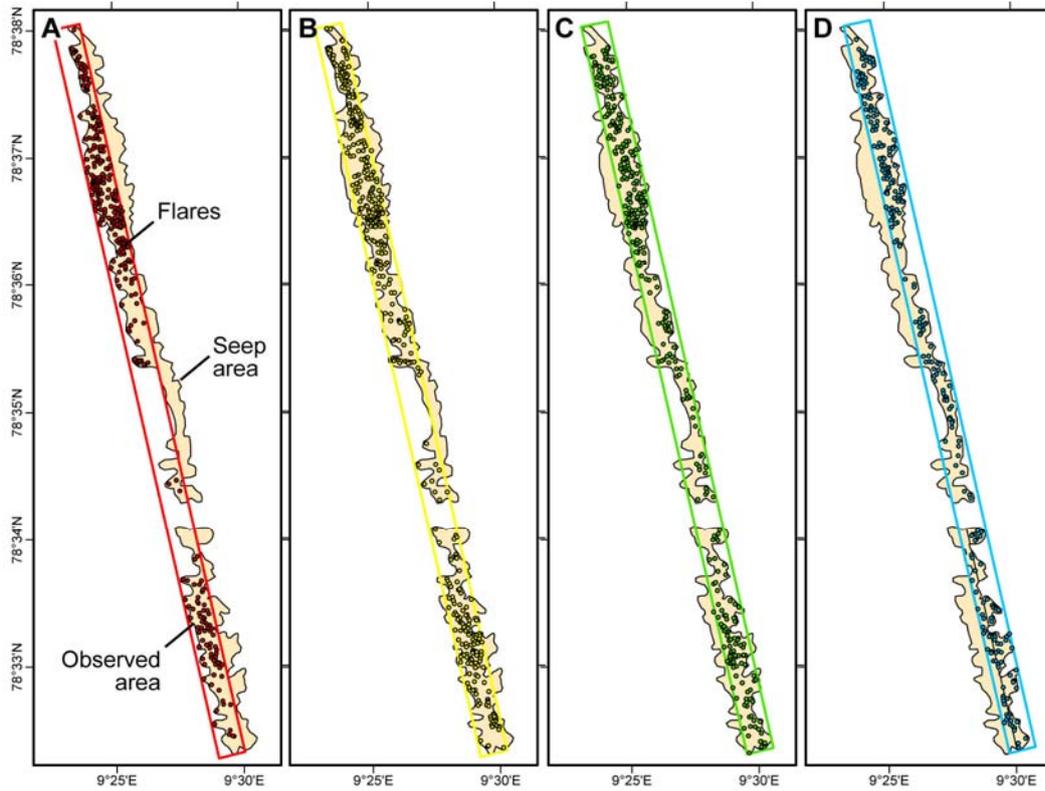
902

903 Figure 7. Screenshot of the record from the horizontally-looking sonar (Sonar Tritech,  
904 625 kHz, 6 m range) mounted on the ROV (Dive 05, 14:37:27 UTC). The image shows  
905 the five bubble streams S1 to S5 at cluster C5 at the western edge of the dive track.



906

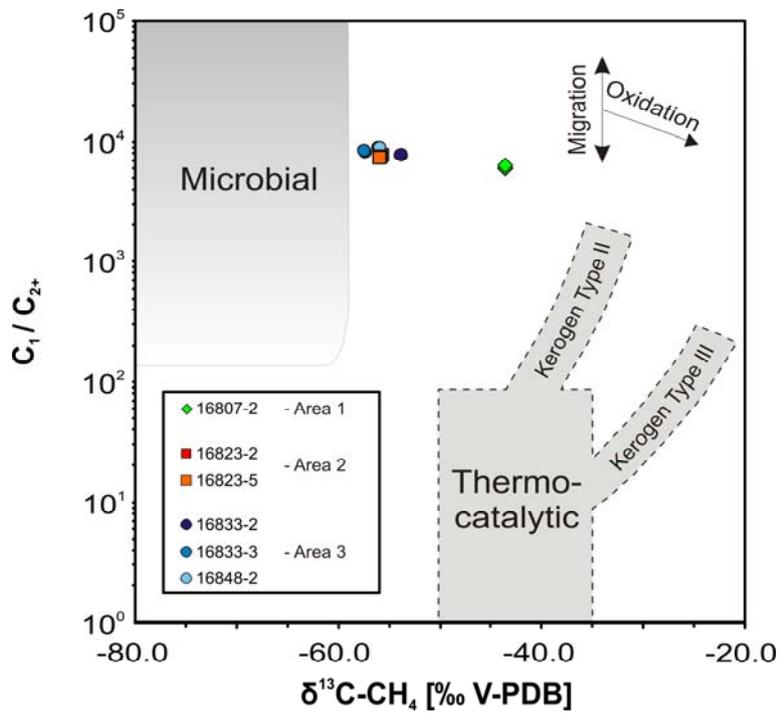
907 Figure 8. Position of flares (circles) found in Area 3 ('396-m flares') plotted on shaded  
 908 bathymetry. Flares were picked in multibeam water column data; the coverage is shown  
 909 as dark grey shading around the ship track (lines). In this study we defined a 'seep area'  
 910 (light grey shading) in which the number of flares was quantified using the four central  
 911 profiles (see Figure 9). The approximate locations where the three ROV dives (06-08) are  
 912 indicated.



913

914 Figure 9. Diagram illustrating the approach of quantifying flares in Area 3. Numbers of  
 915 flares were estimated within the ‘observed area’, which is the region covered by  
 916 multibeam (rectangular box) intersecting with the ‘seep area’. Occurrence of flares were  
 917 repeatedly determined along four parallel profiles (A-D) as indicated in Figure 8 and  
 918 summarized in Table 4.

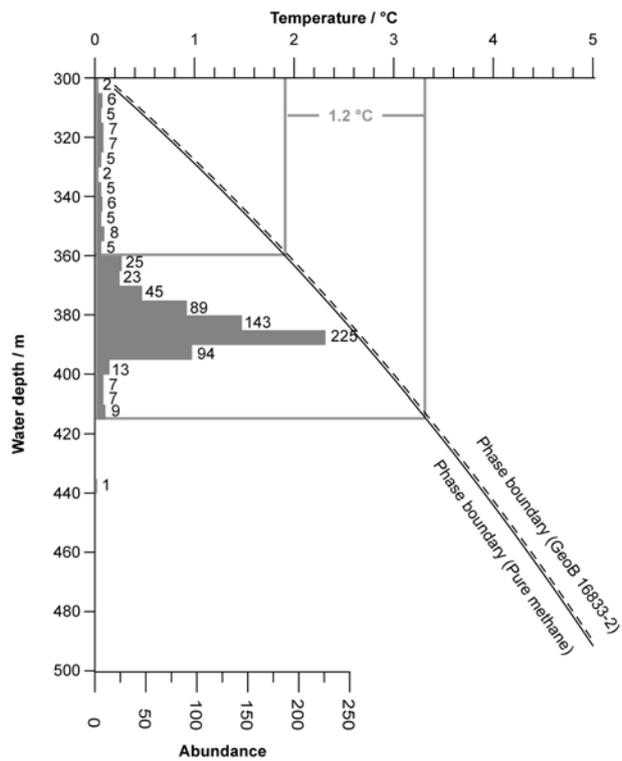
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920

921 Figure 10. Molecular ( $\text{C}_1/\text{C}_{2+}$ ) vs. stable C isotopic composition of methane ( $\delta^{13}\text{C-CH}_4$ )  
922 sampled in Areas 1–3. Classification according to the ‘Bernard diagram’ modified after  
923 Whiticar (1990). Gas samples studied herein plot to close to the empirical field of  
924 microbial methane except for those from Area 1.

925



926

927 Figure 11. Composite figure showing the hydrate (structure I) phase boundary and the  
928 abundance of flares in Area 3 in 5m-depth intervals. Phase boundaries were calculated  
929 considering bottom water salinity and the molecular composition of (i) gas sample GeoB  
930 16833-2 collected with the Gas Bubble Sampler in Area 3 (Table 2) and (ii) pure  
931 methane.